



InSight

JPL



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ETH

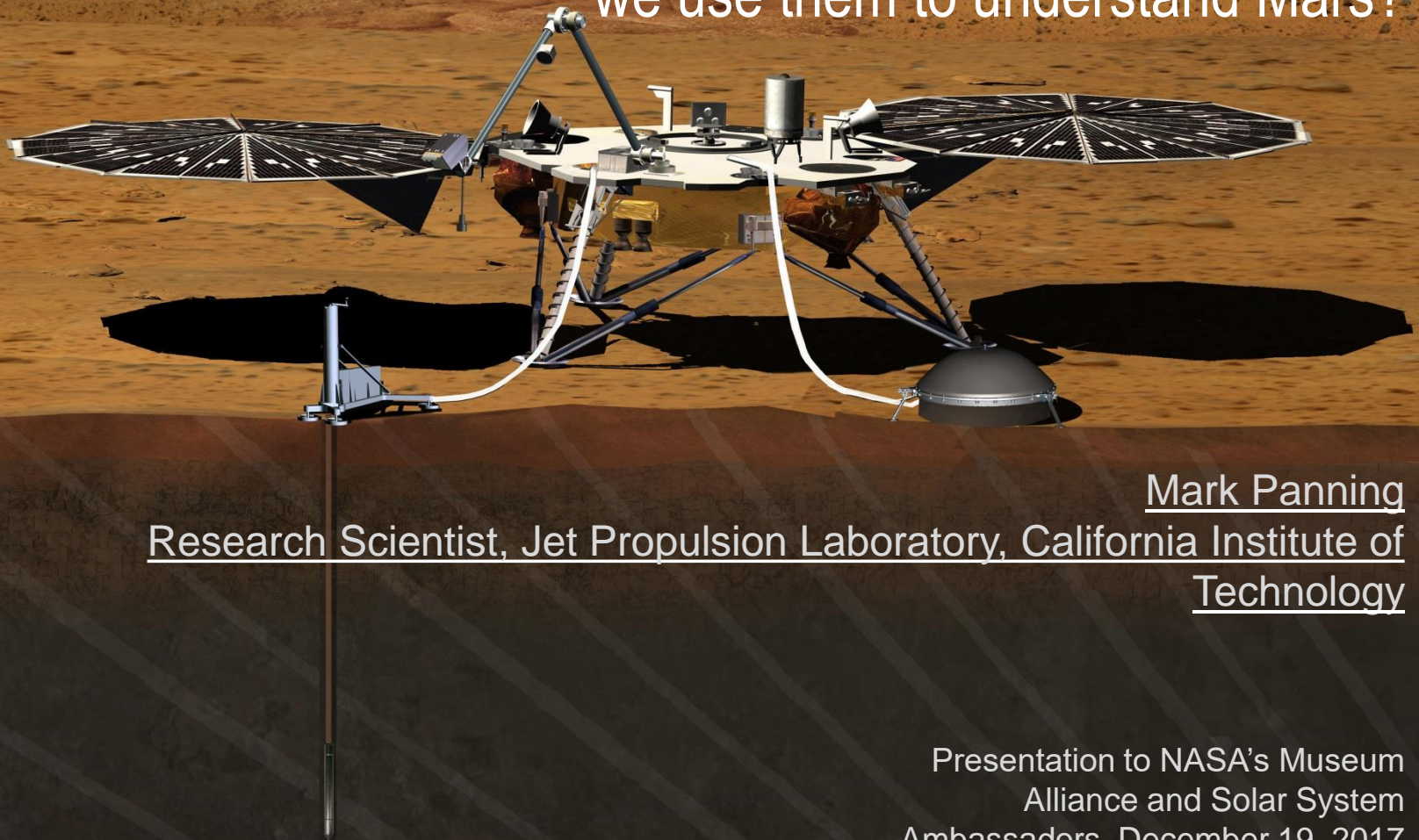


Imperial College
London



InSight

Marsquakes: How do we detect them, and how do we use them to understand Mars?



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Technology

Presentation to NASA's Museum
Alliance and Solar System
Ambassadors, December 19, 2017

A more basic question: What are earthquakes?

Obviously, the biggest reason most people think about earthquakes is in the context of natural hazards, but seismologists use them to learn about the physics of what cause them, and what the Earth looks like between where the quakes happen and where we observe them

San Francisco, 1906



Earthquakes (and marsquakes) happen on faults





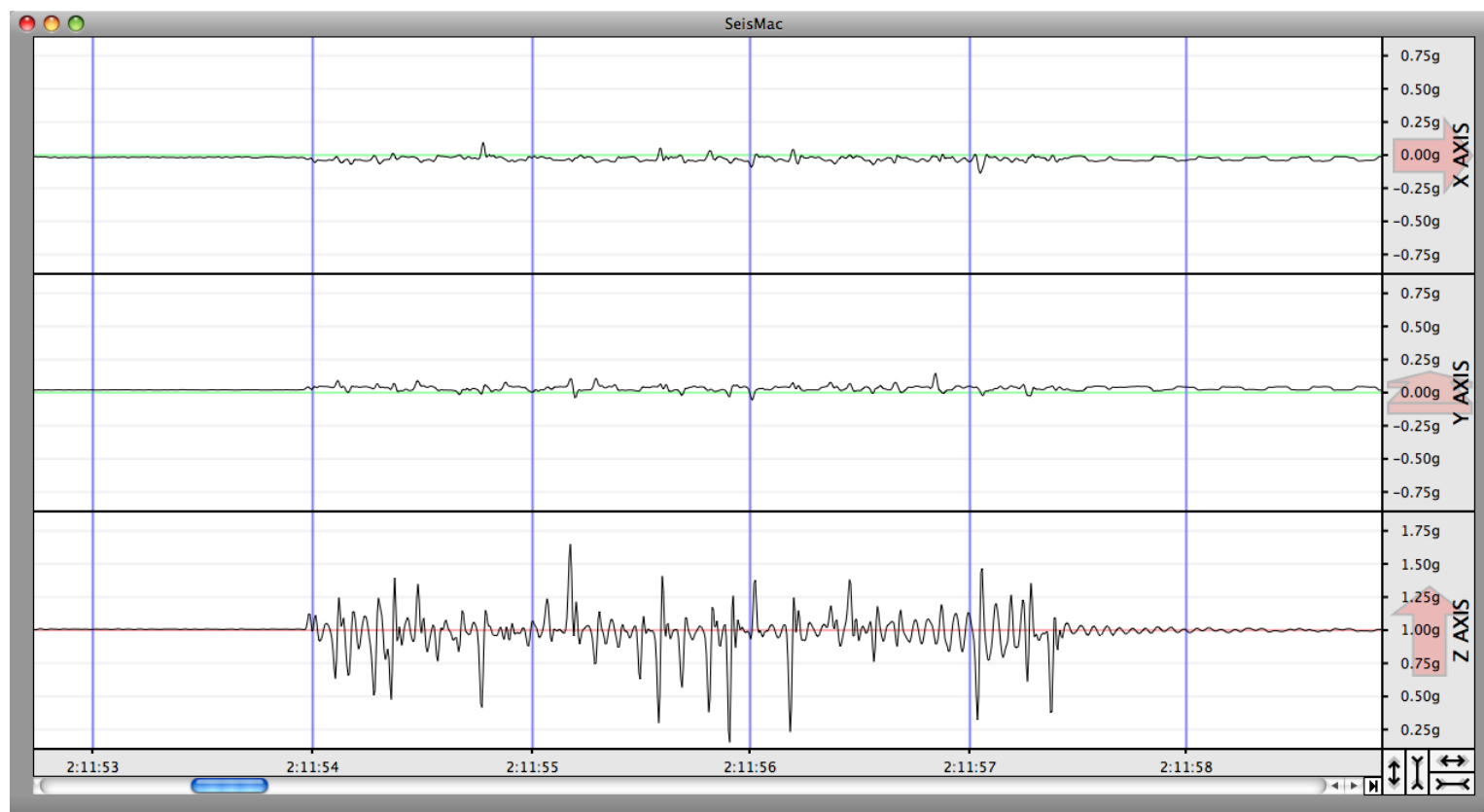
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Quakes make waves



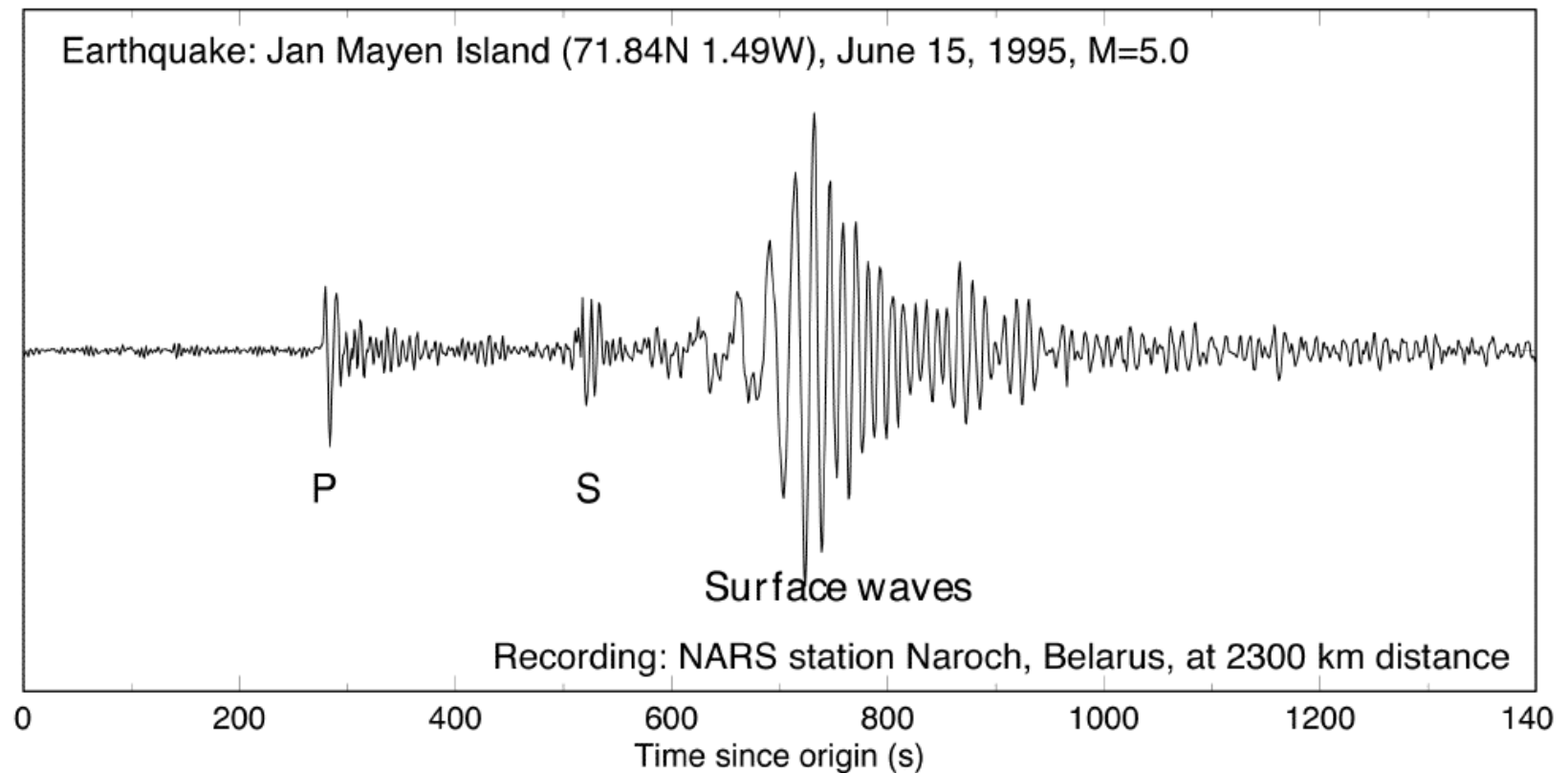
Marsquake visualization from Martin van Driel, ETH Zürich

Example output from SeisMac, which makes seismograms from the seismometer inside a laptop (most laptops and phones have accelerometers, compared with a seismogram from an earthquake)



Seismometers measure waves

Example output from SeisMac, which makes seismograms from the seismometer inside a laptop (most laptops and phones have accelerometers, compared with a seismogram from an earthquake)



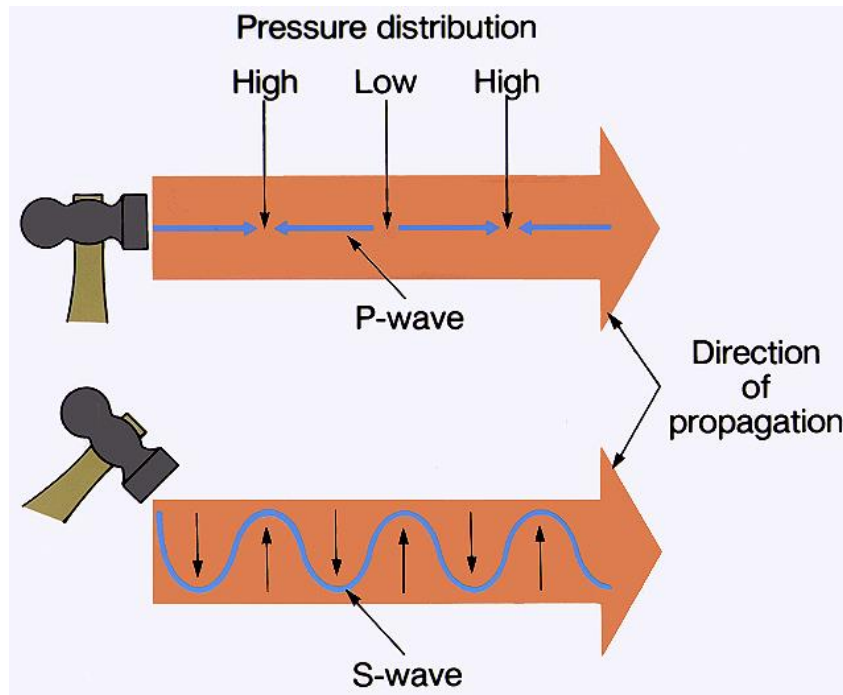
EQ faulting: ~1 second

Seismogram: ~12 minutes

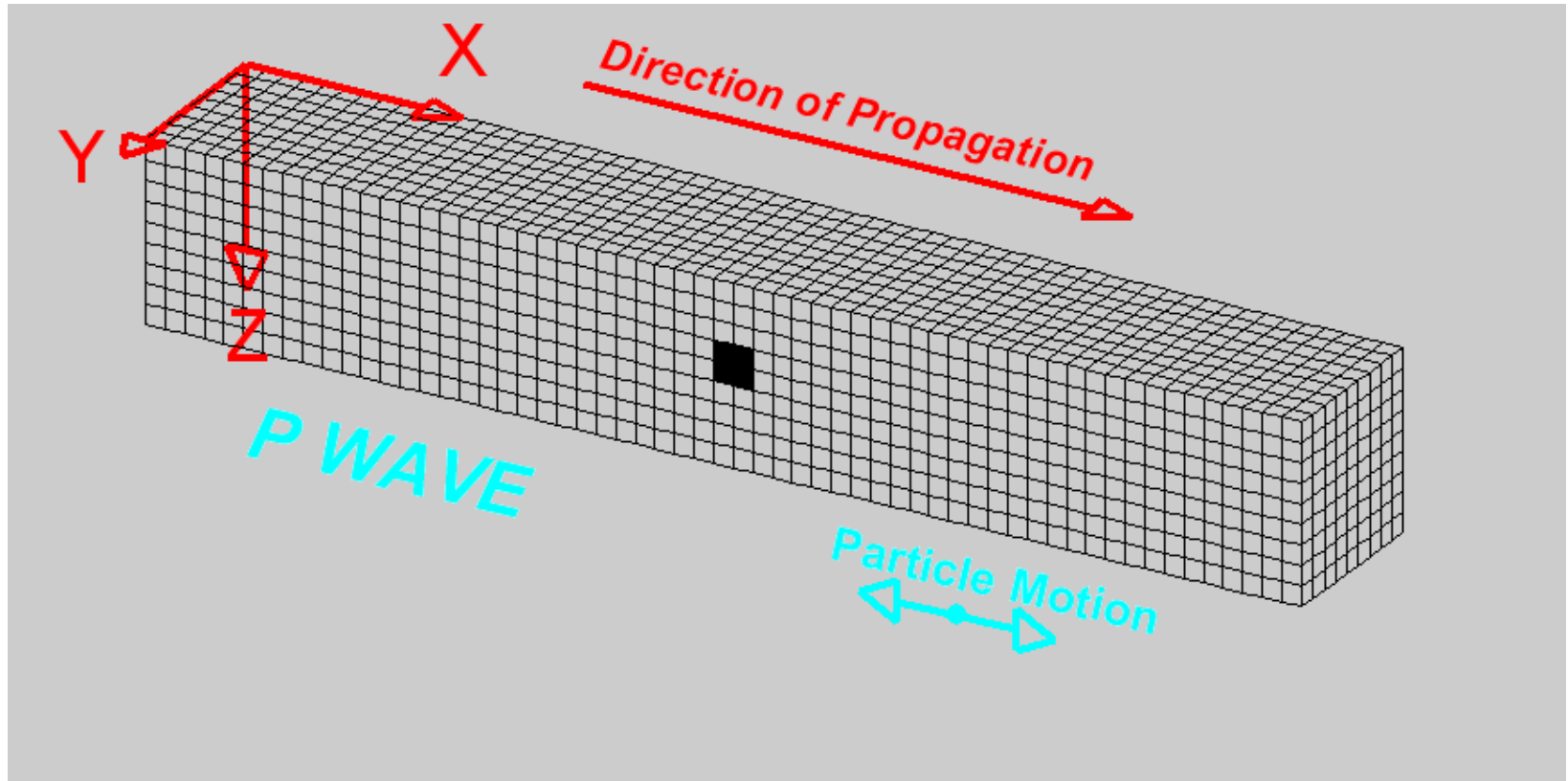


Wherever they are, quakes make the same kind of waves

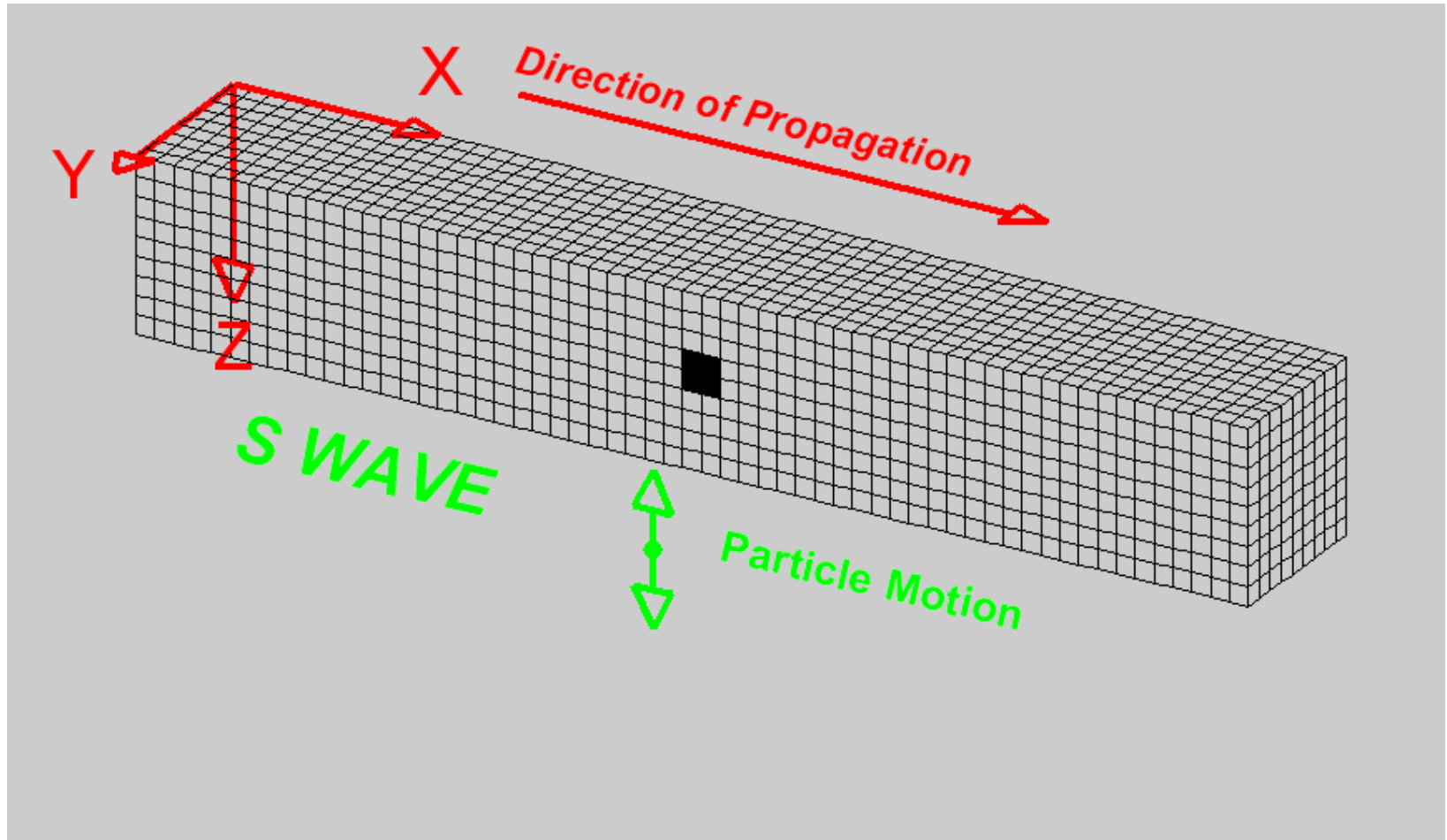
Body waves go through the interior (or body) of the planet, and can be divided into two types:



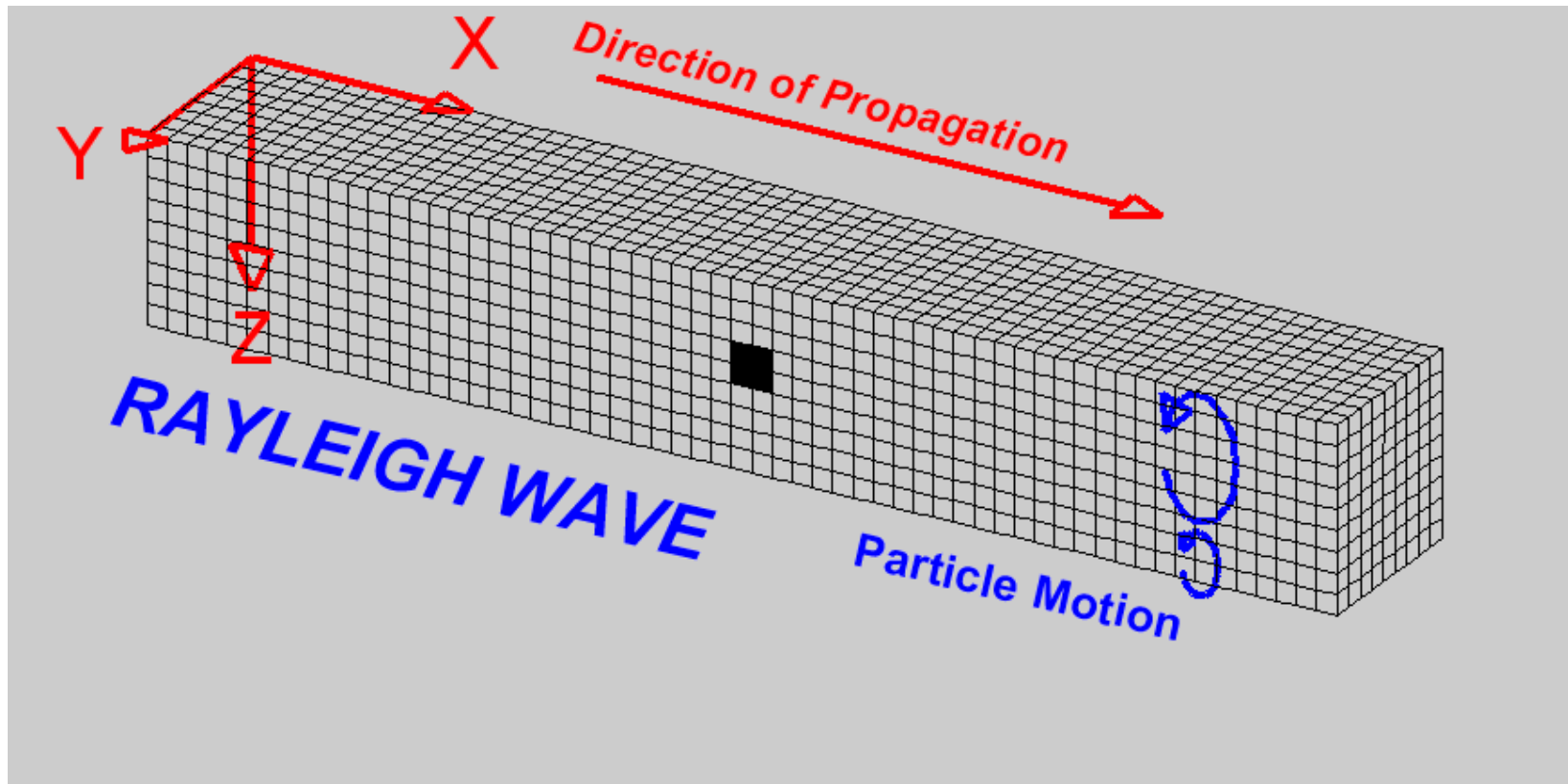
Surface waves travel around stuck to the top of the planet



<http://web.ics.purdue.edu/~braile/edumod/waves/WaveDemo.htm>



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- Planetary missions tend to focus on surface observations, which do not tell us directly about what's inside planets
- Most of our detailed knowledge of the interior structure of the Earth comes from seismology
- Most of what we know about planetary interiors comes from gravity measurements plus educated guesses about chemistry

Comparing Earth, Moon, and Mars on the same scale

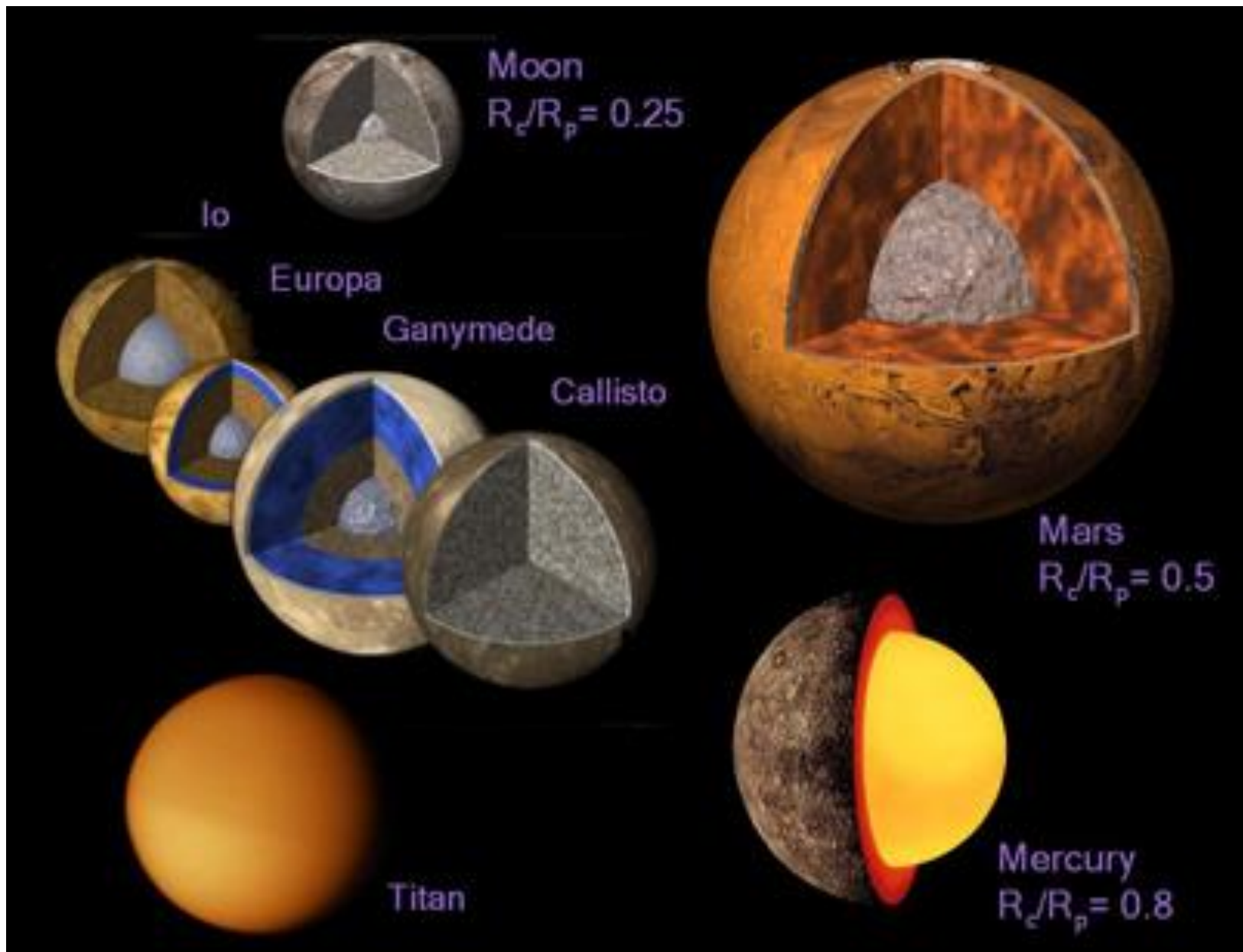


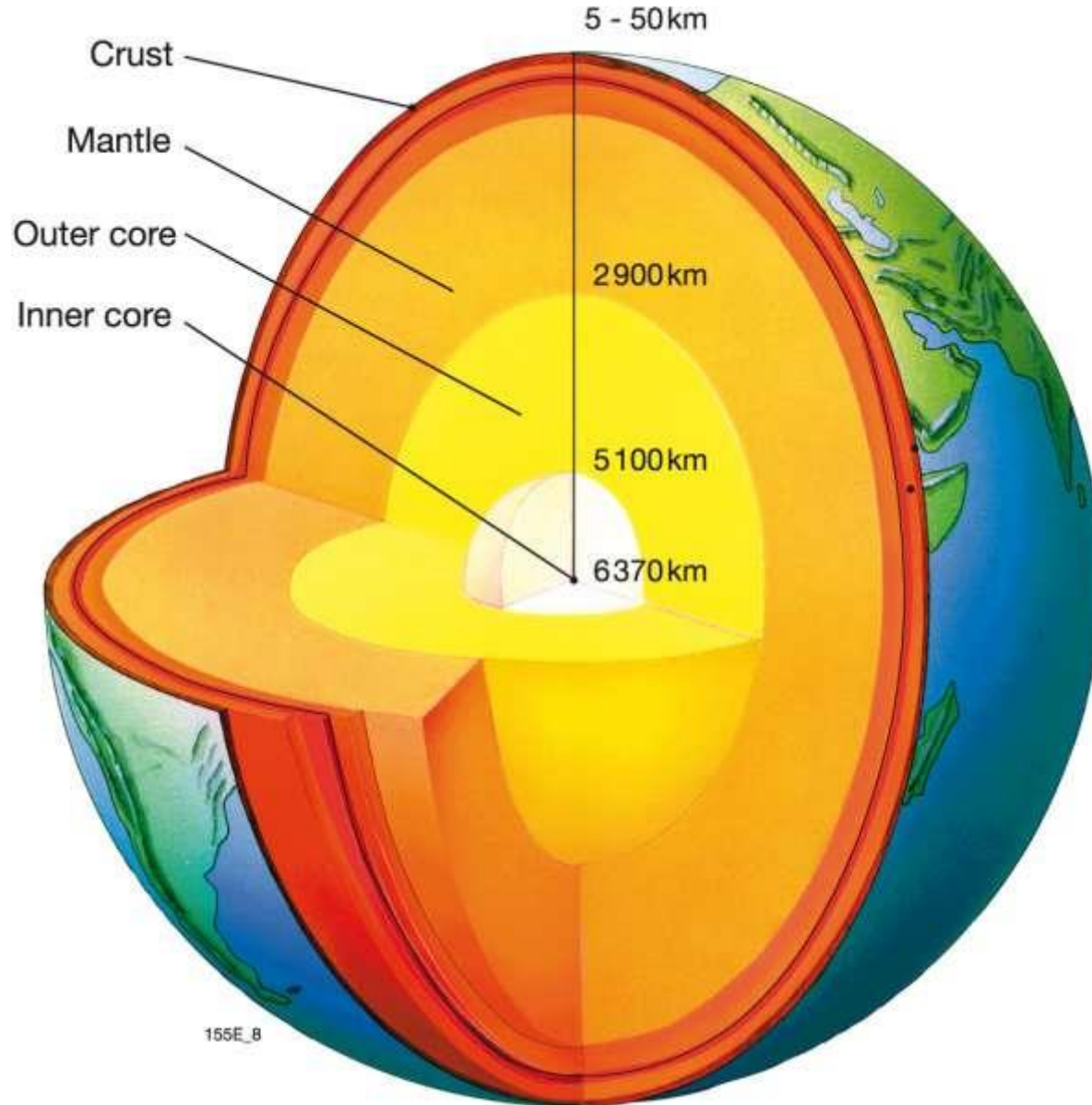
Earth

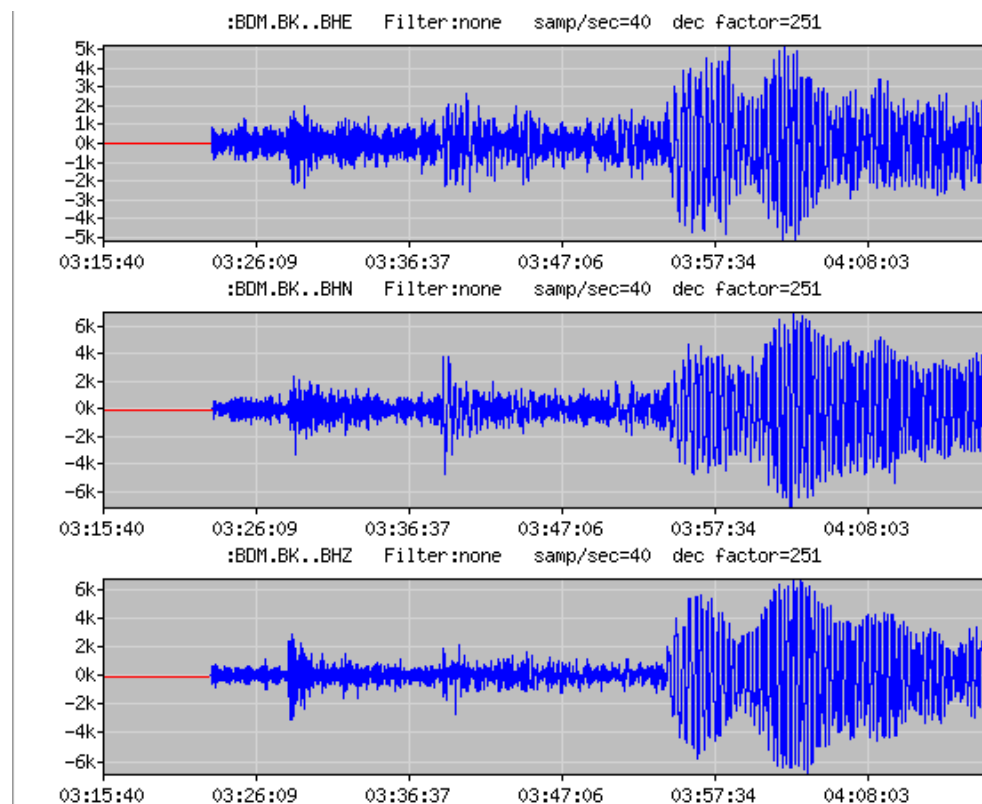
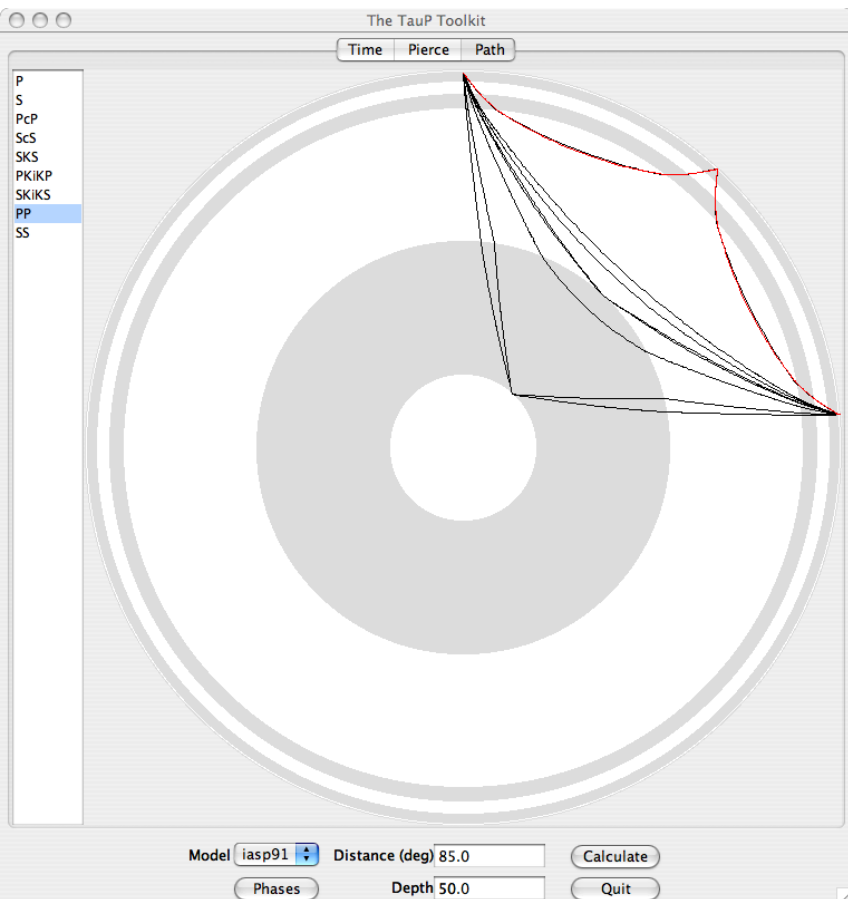
Mars

Moon

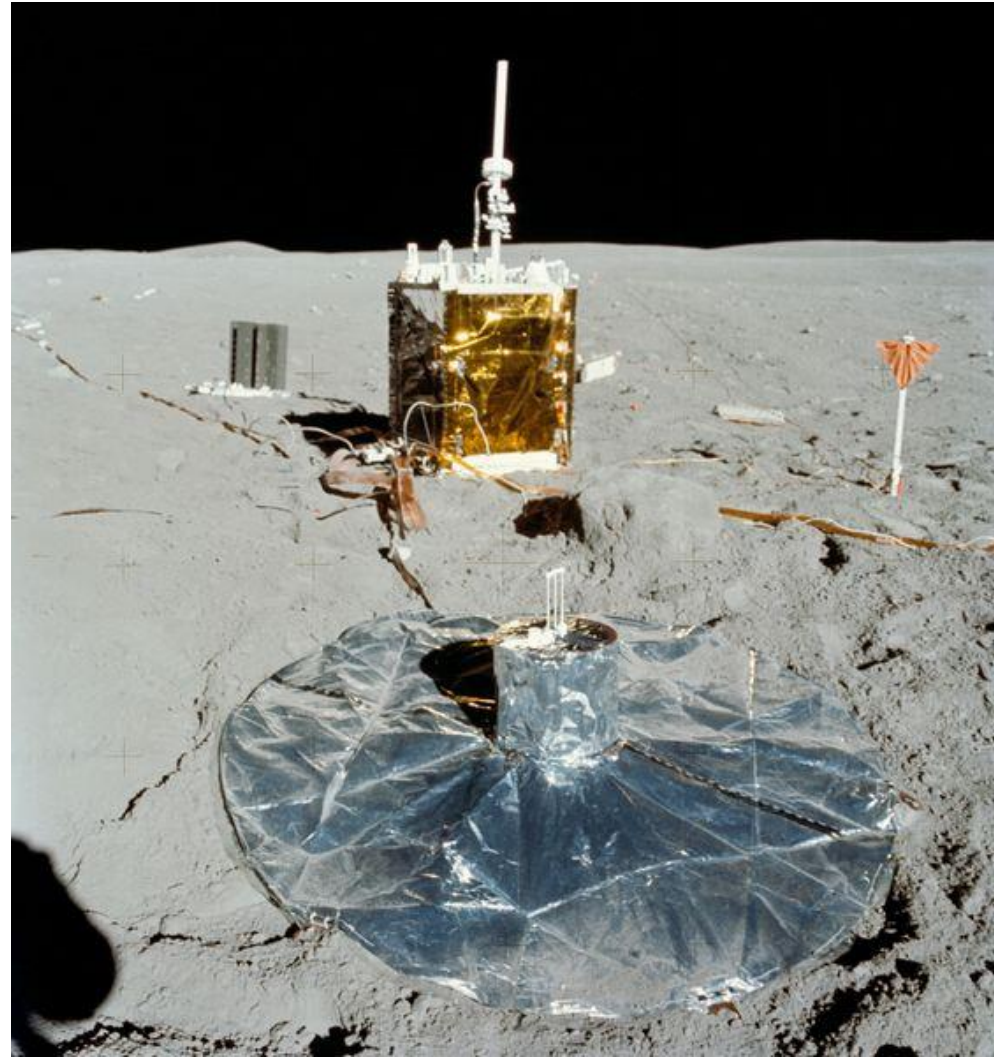
Planetary interiors, or at least our best guesses





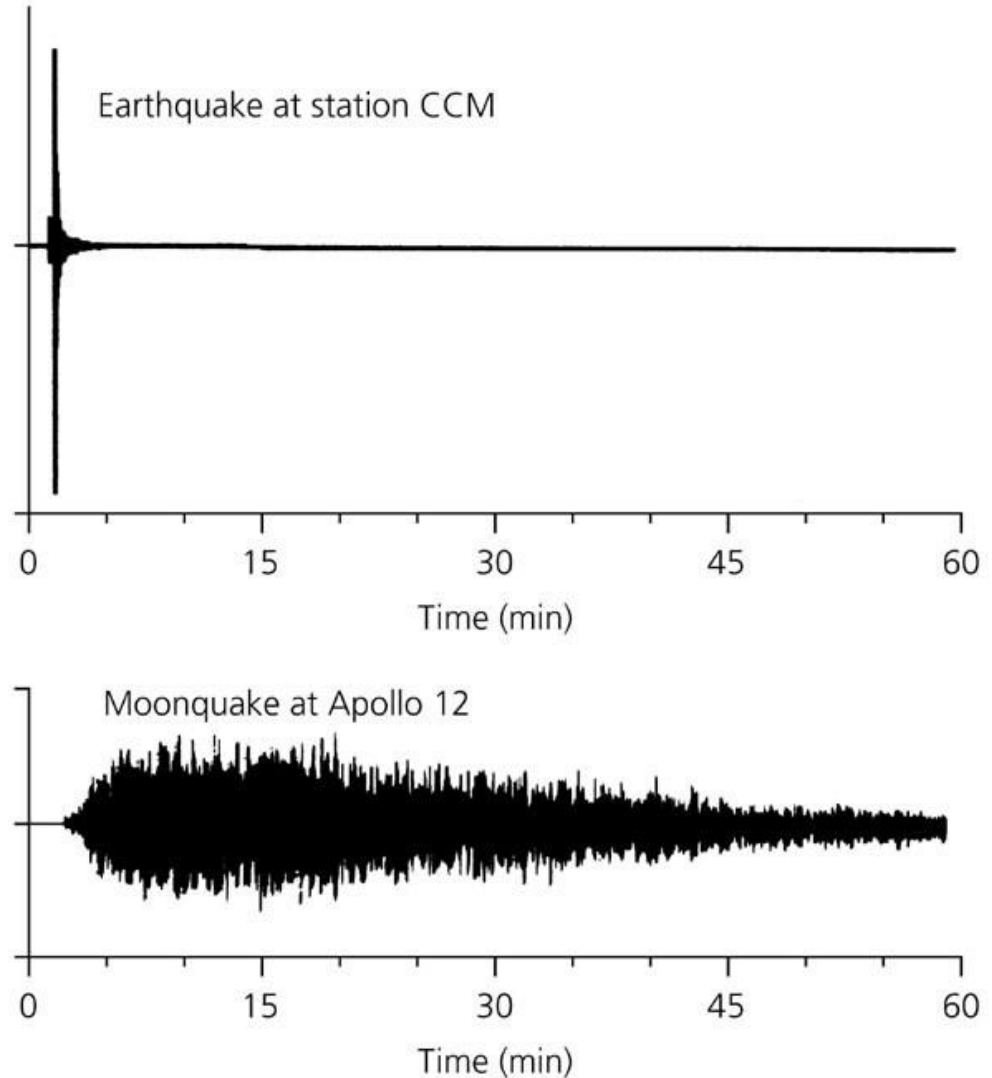


- The “good”: 5 seismometers were placed on Earth’s moon and recorded data (Apollo 11, 12, 14, 15, and 16 missions 1969-1972)
- Data recorded until 1977



- The “good”: 5 seismometers were placed on Earth’s moon and recorded data (Apollo 11, 12, 14, 15, and 16 missions 1969-1972)
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Figure 3.7-10: Comparison of seismograms on the earth and moon.



- The “bad”:
- Viking landers included seismometers
- 1 didn’t “uncage” – no data
- The other primarily recorded the rocking of the lander due to wind

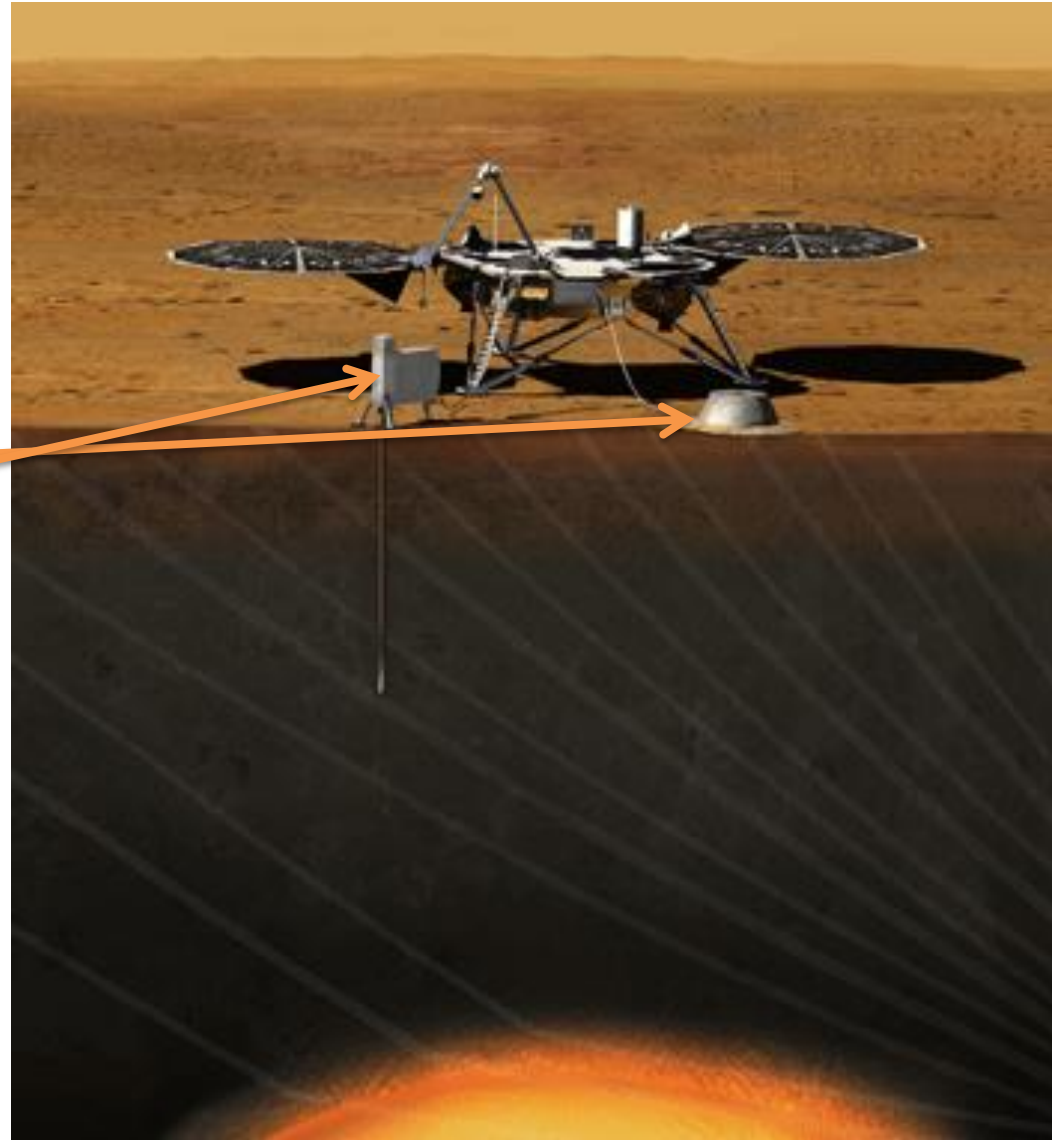


- The “ugly”: At least 10 other seismometers have been included in launched missions that failed for a variety of reasons



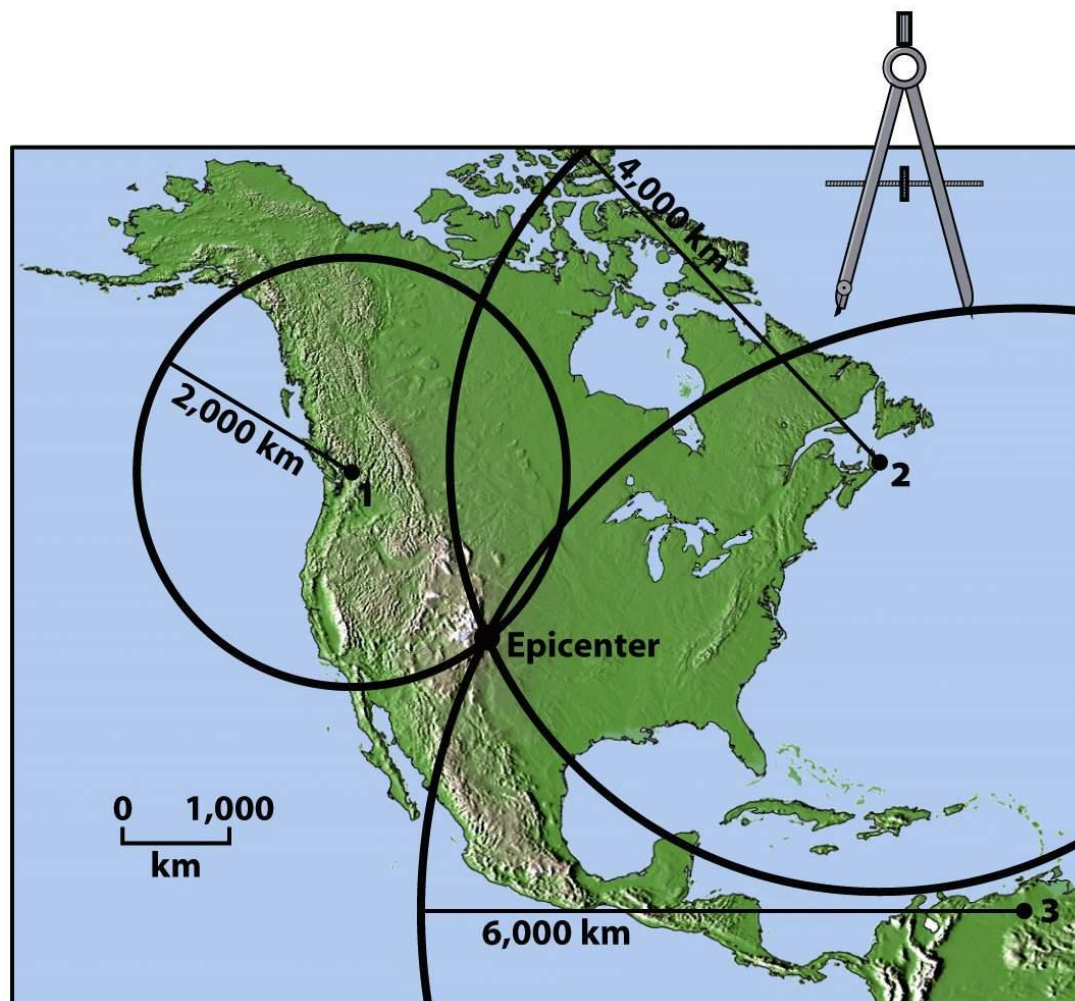
Mars 96 – failed due to launch vehicle stage failure – never escaped Earth orbit

- Launches May 5, 2018 (6 months travel time)
- Instruments on board
 - SEIS
 - HP³
 - RISE
 - Cameras
- Should return data for ~1 martian year (about 2 Earth years)



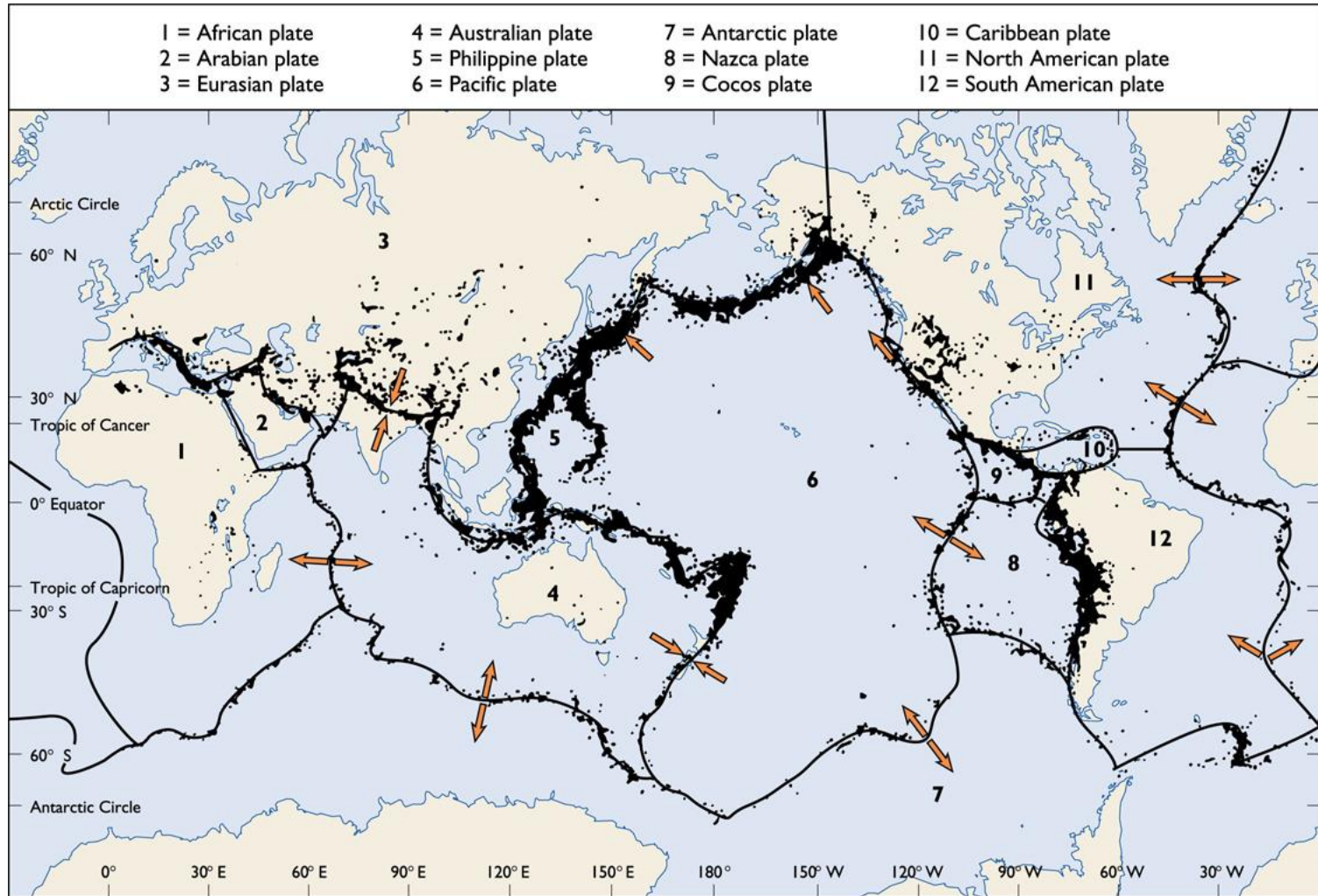


- Modern seismology relies on networks
 - Accurate locations and origin times
 - Array seismology techniques

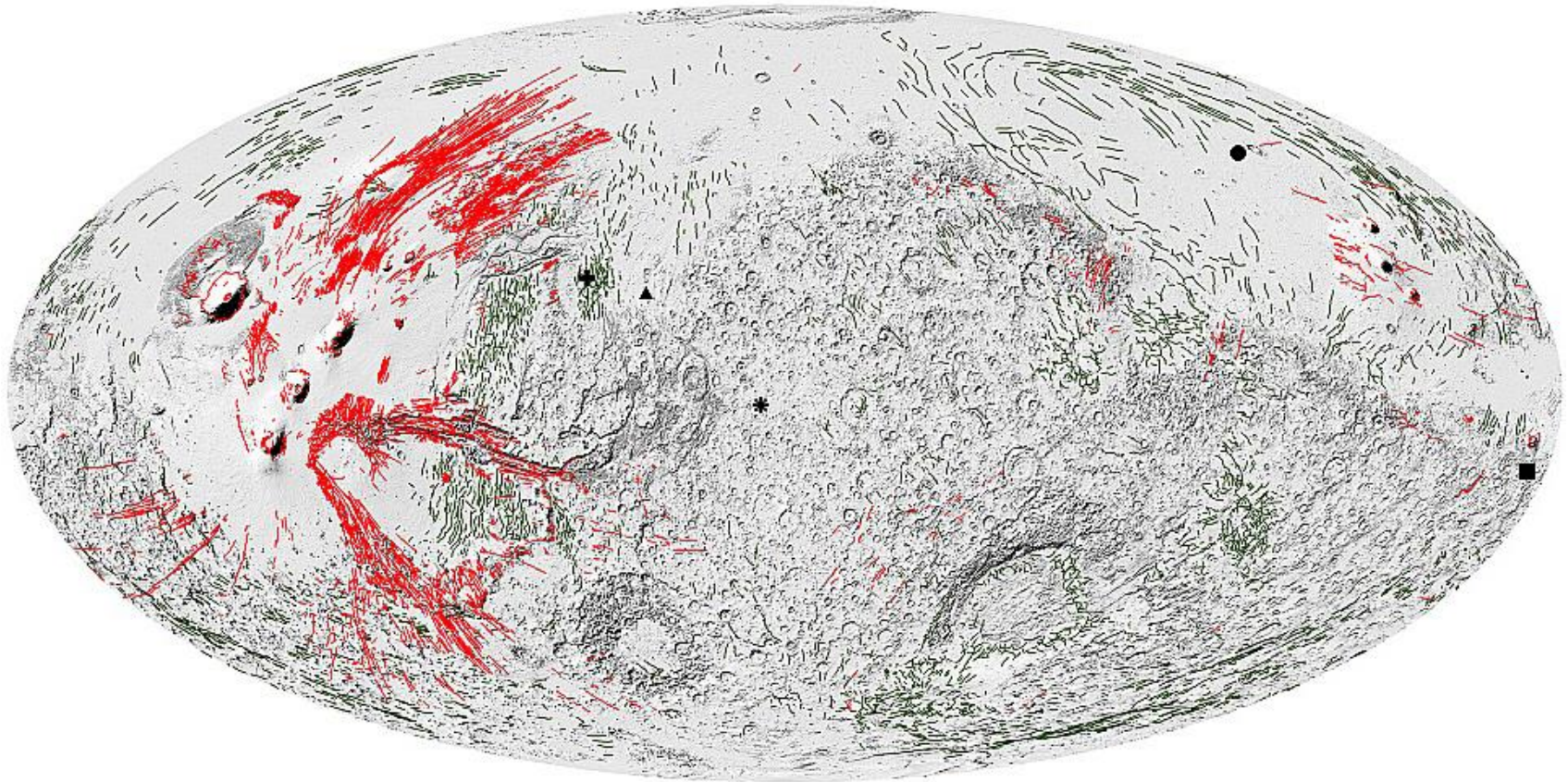


- Having a network on Mars would be wonderful (and it's been proposed many times)
- Having one station allows for an infinite relative increase in our seismic data, though!
- Two options for how to proceed if we want to model the interior seismic structure of Mars
 - Techniques that don't require location information
 - Single station location techniques

Where do we see earthquakes?



Earthquakes, shown as black dots, occur at boundaries of large plates

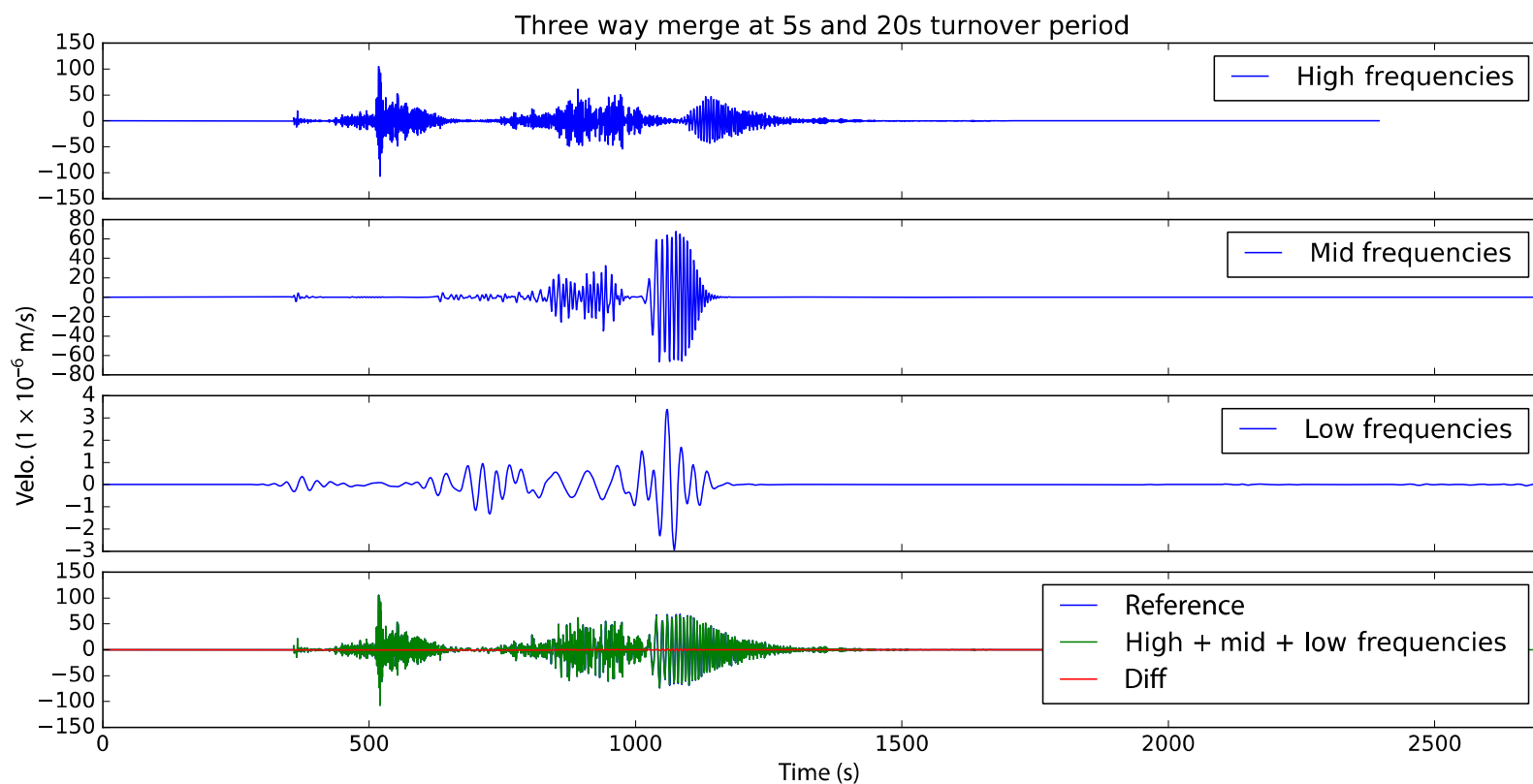


Fault map from Martin Knapmeyer, showing faults observed from orbital images in red and green (no plate tectonics on Mars, but plenty of faults)

- Viking Lander II
- ✚ Viking Lander I
- ▲ Mars Pathfinder
- Spirit
- * Opportunity

What will seismograms look like on Mars?

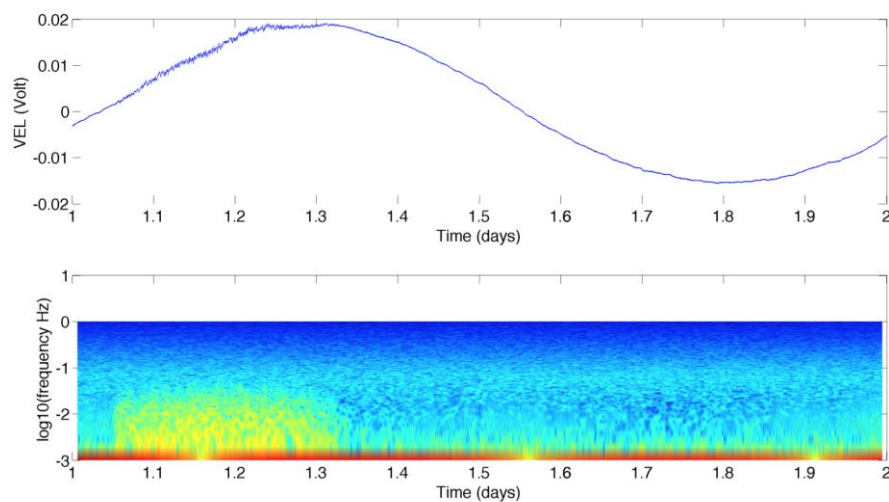
- ✓ We can use computational techniques to simulate seismograms through Mars



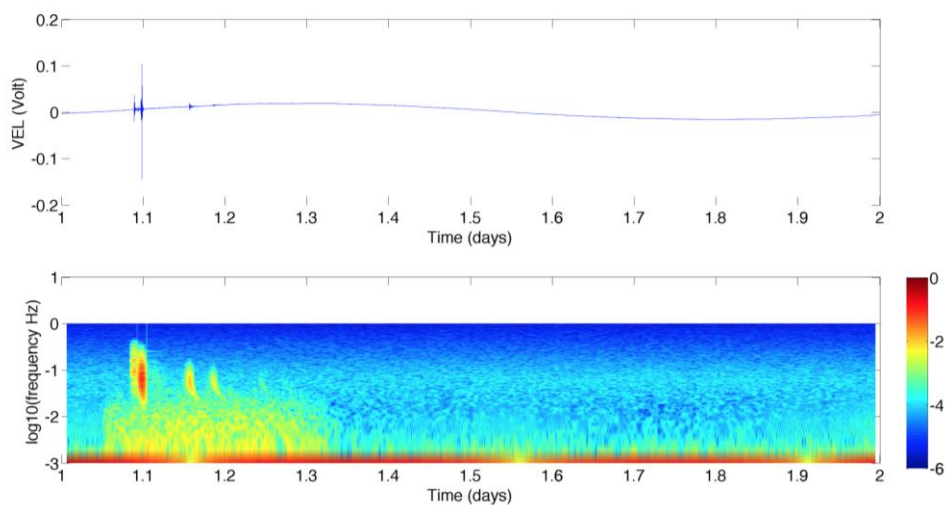
What will seismograms look like on Mars?

- ✓ We can add simulated noise from atmospheric, temperature, and even tidal variations

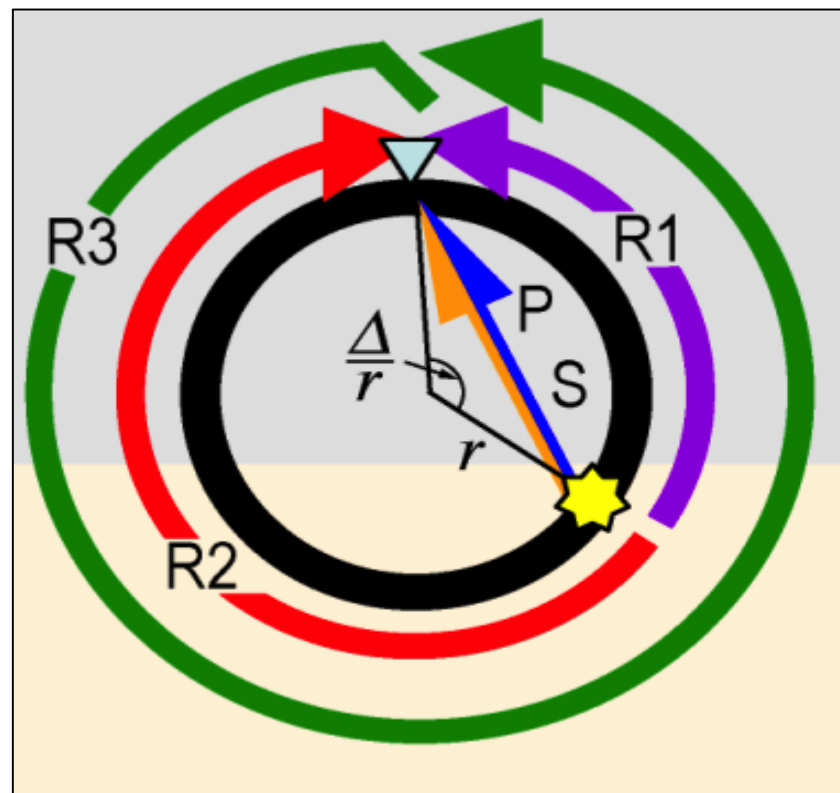
Noise

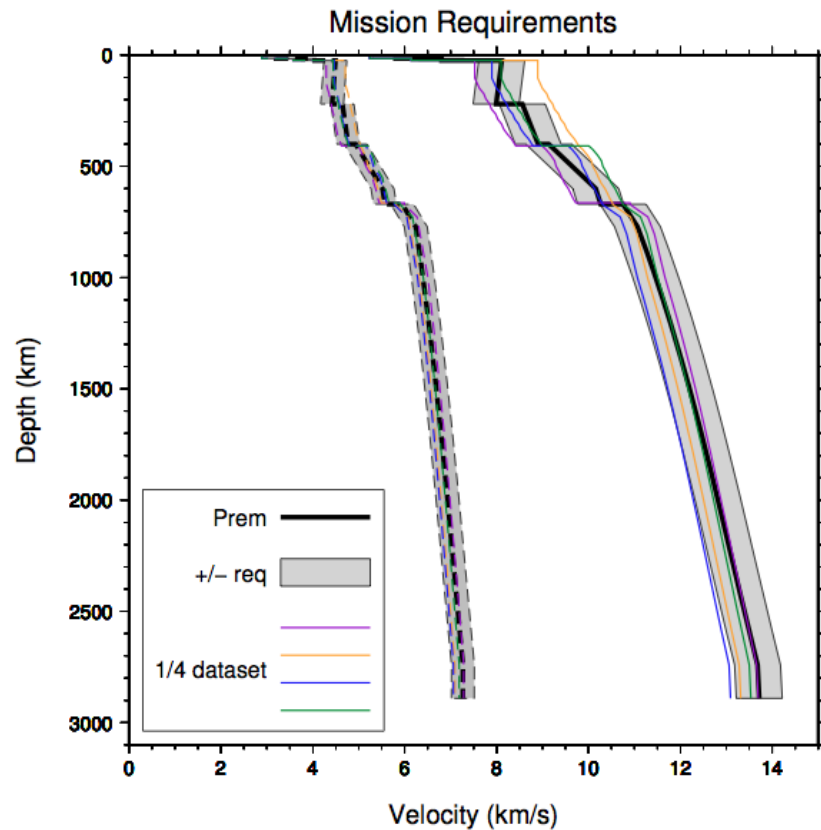


Marsquake + noise

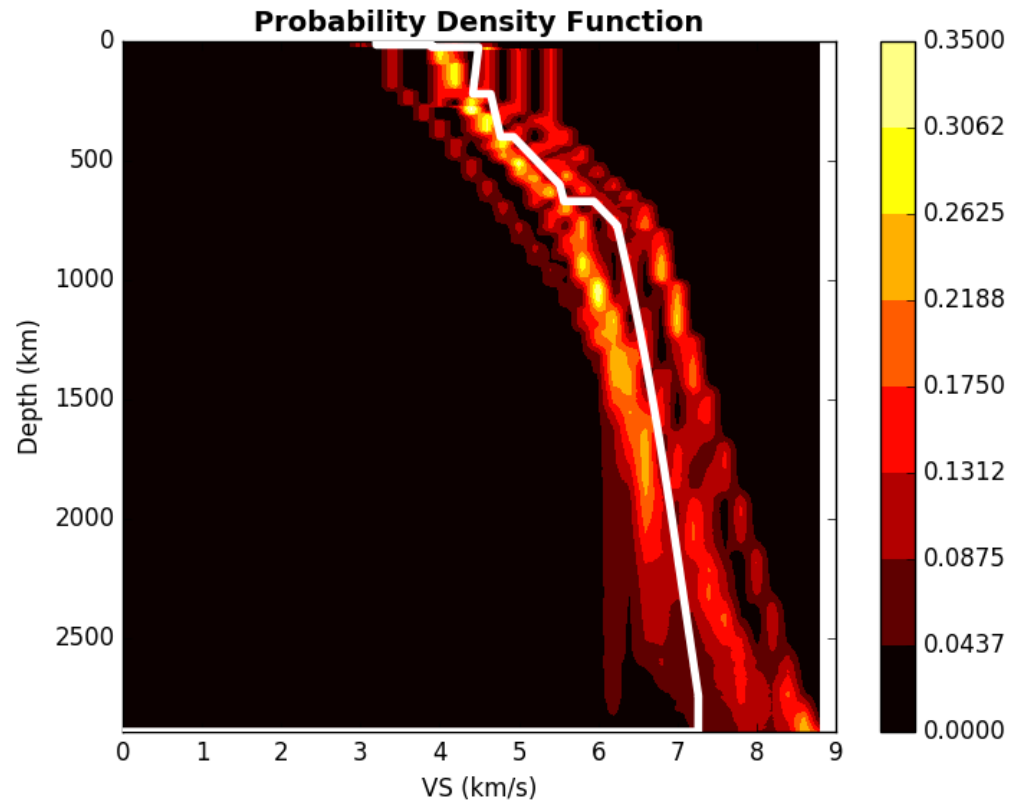


- We estimate location using the timing of 3 orbits of surface waves (called R1, R2, and R3). There are 3 unknowns: Δ , the distance, t_0 , the origin time, and U , the surface wave group velocity.
 - $U = 2\pi / (R3 - R1)$
 - $\Delta = \pi - U(R2 - R1) / 2$
 - $t_0 = R1 - \Delta / U$





With 7 “big” events



With only 5 smaller events



- To know about the inside of a planet, nothing's better than seismology
- We're going to land a seismic station on Mars next year
- We should see marsquakes, and we should be able to use them to figure out the internal details of Mars